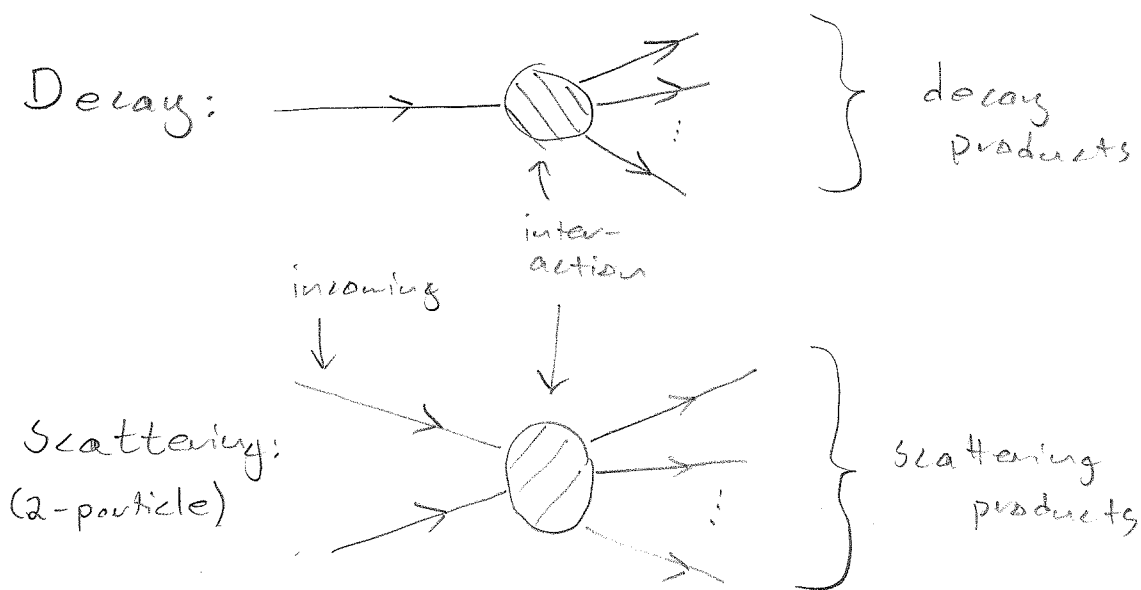


5. Decays

In previous chapters we looked at solutions for free particles. Let us now switch on interactions. This leads to 2 processes:



• Decays:

- lifetime $\tau = \frac{1}{\Gamma}$, Γ decay rate ($N(t) = N_0 e^{-\Gamma t}$)

- decay channel: which particles produced

e.g. $\tau^- \rightarrow \mu^- + \bar{\nu}_\mu + \nu_e$

$\tau^- \rightarrow e^- + \bar{\nu}_e + \nu_\tau$

....

* rate for channel i Γ_i ; $\Gamma_{\text{tot}} = \sum_i \Gamma_i$

* branching ratio $\equiv \Gamma_i / \Gamma_{\text{tot}}$

= fraction of decays to channel i

How to calculate the rate?

In general, we need to calculate so-called scattering matrix (S-matrix)

$$S_{FI} = \langle \phi_F(t_F) | \hat{U}(t_F, t_I) | \phi_I(t_I) \rangle$$

↑
final state

↑ initial state
"time evolution"
operator

take $t_i \rightarrow -\infty$; $t_F \rightarrow +\infty$

Let us see how this happens:

Time evolution - equivalent viewpoints, pictures:

* Schrödinger picture

- States depend on t , operators not

$$i \partial_t |\psi\rangle_S = \hat{H} |\psi\rangle_S$$

$$\Rightarrow |\psi(t)\rangle_S = e^{-i\hat{H}t} |\psi(0)\rangle_S$$

$$- \langle \hat{A}_S \rangle = \langle_S \psi(t) | \hat{A}_S | \psi(t) \rangle_S$$

* Heisenberg picture

- States independent of t , operators depend on t

$$\hat{A}_H \equiv e^{i\hat{H}t} \hat{A}_S e^{-i\hat{H}t}$$

$$\Rightarrow i \frac{\partial}{\partial t} \hat{A}_H = [\hat{A}_H, \hat{H}]$$

$$- \langle \hat{A}_H(t) \rangle = \langle_S \psi(0) | \hat{A}_H(t) | \psi(0) \rangle_S = \langle \hat{A}_S \rangle$$

* Interaction picture (Dirac-picture)

- Assume $\hat{H} = \hat{H}_0 + g\hat{V}$, g small

\hat{H}_0 : free Hamiltonian, \hat{V} : interaction

- Define

$$|\psi(t)\rangle_I = e^{i\hat{H}_0 t} |\psi(t)\rangle_S = e^{i\hat{H}_0 t} e^{-i\hat{H} t} |\psi(0)\rangle_S$$

$$\hat{A}_I(t) = e^{i\hat{H}_0 t} \hat{A}_S e^{-i\hat{H}_0 t}$$

$$\hat{V}_I(t) = e^{i\hat{H}_0 t} \hat{V} e^{-i\hat{H}_0 t}$$

$$\begin{aligned} \Rightarrow i \frac{d}{dt} |\psi(t)\rangle_I &= e^{i\hat{H}_0 t} (-\hat{H}_0 + \hat{H}) |\psi(t)\rangle_I \\ &= g \hat{V}_I(t) |\psi(t)\rangle_I \end{aligned}$$

$$i \frac{d}{dt} \hat{A}_I(t) = [\hat{A}_I(t), \hat{H}_0]$$

(if $g\hat{V}=0$, equivalent to Heisenberg picture)

- Define time evolution operator

$$|\psi(t)\rangle_I = \hat{U}_I(t, t_0) |\psi(t_0)\rangle_I$$

$$\Rightarrow i \frac{d}{dt} \hat{U}_I(t, t_0) = g \hat{V}_I(t) \hat{U}_I(t, t_0) \quad (*)$$

Note: $\hat{U}_I(t_0, t_0) = \mathbb{1}$

* All pictures equivalent - choose suitable one for computations

* Interaction picture \leftrightarrow perturbation theory

Clearly, solution for (*) is

$$\hat{U}_I(t, t_0) = \mathbb{1} - ig \int_{t_0}^t dt' \hat{V}_I(t') \hat{U}_I(t', t_0)$$

This can be iterated to give perturbation series:

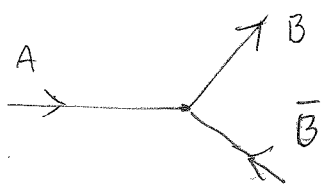
$$\hat{U}_I(t, t_0) = \mathbb{1} - ig \int_{t_0}^t \hat{V}_I(t') dt' + \mathcal{O}(g^2)$$

(Exercise: do this to order g^2)

Let us apply this to decay process.

For concreteness, we take process

where heavy particle A decays into light particle + antiparticle B, \bar{B}



(This could be an effective theory for decay of ρ -meson into π^+, π^- - in QCD the process is much more complicated.)

Interaction:

$$g\hat{V}_I = \int d^3\vec{x} M \underbrace{\hat{\phi}_B(\vec{x}, t) \hat{\phi}_B^\dagger(\vec{x}, t)}_{\text{constant (here)}} \hat{\phi}_A(\vec{x}, t)$$

Here

$$\hat{\phi}_A(\bar{x}, t) = \int \frac{d^3\bar{p}}{\sqrt{(2\pi)^3 2E_{\bar{p}}}} \left[\hat{a}_{\bar{p}}^{(A)} e^{-ip \cdot x} + \hat{a}_{\bar{p}}^{+(A)} e^{ip \cdot x} \right] \in \mathbb{R}$$

↑ destroys incoming A

$$\hat{\phi}_B(\bar{x}, t) = \int \frac{d^3\bar{p}}{\sqrt{(2\pi)^3 2E_{\bar{p}}}} \left[\hat{a}_{\bar{p}}^{(B)} e^{-ip \cdot x} + \hat{b}_{\bar{p}}^{+(B)} e^{ip \cdot x} \right]$$

↑ creates outgoing B

$$\hat{\phi}_B^+(\bar{x}, t) = \int \frac{d^3\bar{p}}{\sqrt{(2\pi)^3 2E_{\bar{p}}}} \left[\hat{a}_{\bar{p}}^{+(B)} e^{ip \cdot x} + \hat{b}_{\bar{p}}^{(B)} e^{-ip \cdot x} \right]$$

↑ creates outgoing B

Thus, the non-trivial part of S-matrix in 1st order is

$$S_{FI} = -i \langle B(\bar{p}_1) \bar{B}(\bar{p}_2) | \int dt d^3\bar{x} M \hat{\phi}_B^+ \hat{\phi}_B \hat{\phi}_A | A(\bar{q}) \rangle$$

$$= -i \int dt \int d^3\bar{x} M \frac{1}{\sqrt{(2\pi)^3 2E_{\bar{q}}} \sqrt{(2\pi)^3 2E_{\bar{p}_1}} \sqrt{(2\pi)^3 2E_{\bar{p}_2}}} e^{i(p_1 + p_2 - q) \cdot x}$$

$$= -i (2\pi)^4 \delta^{(4)}(p_1 + p_2 - q) \underbrace{\frac{M}{\sqrt{(2\pi)^3 2E_{\bar{q}}} \sqrt{(2\pi)^3 2E_{\bar{p}_1}} \sqrt{(2\pi)^3 2E_{\bar{p}_2}}}}_{\text{Amplitude}}$$

4-momentum conservation

comes from here
(energy + momentum)

Transfer matrix element T_{FI}

(Here $|A(\bar{q})\rangle = \hat{a}_{\bar{q}}^{+(A)} |0\rangle$; $\langle B(\bar{p}_1) \bar{B}(\bar{p}_2) | = \langle 0 | \hat{a}_{\bar{p}_1}^{(B)} \hat{b}_{\bar{p}_2}^{(B)}$
 $= (\hat{a}_{\bar{p}_1}^{(B)+} \hat{b}_{\bar{p}_2}^{(B)+} |0\rangle)^\dagger$)

From the S -matrix we get the probability by squaring:

$$p \sim |\langle F | U | I \rangle|^2 \sim |S_{FI}|^2$$

$$|S_{FI}|^2 = (2\pi)^4 \delta^{(4)}(p_1 + p_2 - q) (2\pi)^4 \delta^{(4)}(0) \frac{|M|^2}{(2\pi)^3 2E_{\vec{p}_1} (2\pi)^3 2E_{\vec{p}_2} (2\pi)^3 2E_{\vec{q}}}$$

Double δ -function: assume again finite V and time T ;

then $\delta^{(4)}(0) \hat{=} \frac{V \cdot T}{(2\pi)^4} \leftarrow 4\text{-volume}$

\Rightarrow we obtain decay rate $\Gamma = \frac{|S_{FI}|^2}{T}$

by dividing out T .

- Above was actually $\Gamma(\bar{q} \rightarrow \bar{p}_1, \bar{p}_2)$. To get total rate, we should integrate over possible final state \bar{p} 's:

$$\Gamma = \int d^3\bar{p}_1 d^3\bar{p}_2 (2\pi)^4 \delta^{(4)}(p_1 + p_2 - q) \frac{V}{(2\pi)^3 2E_{\vec{q}} (2\pi)^3 2E_{\vec{p}_1} (2\pi)^3 2E_{\vec{p}_2}} |M|^2$$

- Still one problem: incoming state was normalised to infinity; $\langle A(\vec{q}) | A(\vec{q}) \rangle = \delta^{(3)}(0)$.
 $\hat{=} \frac{V}{(2\pi)^3}$

To normalise to single particle, we demand $\langle A | A \rangle = 1$, and divide by

$$V/(2\pi)^3$$

Thus, we obtain

$$\Gamma_{A \rightarrow B\bar{B}} = \frac{1}{2 E_{\bar{q}}} \int \frac{d^3 \bar{p}_1}{(2\pi)^3 2 E_{\bar{p}_1}} \int \frac{d^3 \bar{p}_2}{(2\pi)^3 2 E_{\bar{p}_2}} (2\pi)^4 \delta^{(4)}(p_1 + p_2 - q) |M|^2$$

Fermi's golden rule for decays

For n -particle decays

$$\Gamma_{A \rightarrow 1,2,3,\dots,n} = \frac{1}{2 E_{\bar{q}}} \cdot C \cdot \underbrace{\left\{ \prod_{i=1}^n \frac{d^3 \bar{p}_i}{(2\pi)^3 2 E_{\bar{p}_i}^{(i)}} \right\}}_{\text{Phase space integration}} (2\pi)^4 \delta^{(4)}(\sum_i p_i - q) |M|^2$$

Statistical factor: if j identical particles are produced, $C = 1/j!$

In general, M can have \bar{p}_i, \bar{q} -dependence (not in our example)

These results are valid for general 2-body (n -body) decays, not only to $A \rightarrow B\bar{B}$. (subs. $E_{\bar{p}_i} \rightarrow E_{p_i}^{(i)} = \sqrt{m_i^2 + p_i^2}$)

The form above is the most general one.

Depending on amplitude M , \bar{p}_1, \bar{p}_2 -integrations can be (at least partially) performed.

- Phase space integration example: decay $A \rightarrow 1+2$, and assume M constant: Choose rest frame of A ; $\vec{q} = 0$

$$\begin{aligned} \Gamma_{A \rightarrow 1,2} &= \frac{|M|^2}{2m_A} \int \frac{d^3 \vec{p}_1}{(2\pi)^3 2E_{\vec{p}_1}^{(1)}} \int \frac{d^3 \vec{p}_2}{(2\pi)^3 2E_{\vec{p}_2}^{(2)}} (2\pi)^4 \delta(E_{\vec{p}_1} + E_{\vec{p}_2} - m_A) \delta^{(3)}(\vec{p}_1 + \vec{p}_2) \\ &= \frac{|M|^2}{2m_A} \int \frac{d^3 \vec{p}_1}{(2\pi)^2 4} \frac{\delta(m_A - \sqrt{m_1^2 + \vec{p}_1^2} - \sqrt{m_2^2 + \vec{p}_1^2})}{\sqrt{m_1^2 + \vec{p}_1^2} \sqrt{m_2^2 + \vec{p}_1^2}} \end{aligned}$$

Now $\int d^3 \vec{p}_1 \rightarrow 4\pi \int_0^\infty dp p^2$

Let now $E = \sqrt{m_1^2 + p^2} + \sqrt{m_2^2 + p^2}$

$$\Rightarrow dE = \left[\frac{1}{\sqrt{m_1^2 + p^2}} + \frac{1}{\sqrt{m_2^2 + p^2}} \right] p dp$$

$$= \frac{E}{\sqrt{m_1^2 + p^2} \sqrt{m_2^2 + p^2}} p dp$$

$$\begin{aligned} \Rightarrow \Gamma_{A \rightarrow 1,2} &= \frac{|M|^2}{8\pi m_A} \int_{m_1+m_2}^\infty dE \rho(E) \cdot \frac{1}{E} \delta(m_A - E) \\ &= \frac{|M|^2}{8\pi m_A^2} \rho(m_A) \quad , \quad m_1 + m_2 < m_A \end{aligned}$$

where $\rho(m_A)$ can be solved from

$$m_A^2 = m_1^2 + m_2^2 + 2p^2 + 2\sqrt{(m_1^2 + p^2)(m_2^2 + p^2)}$$

$$\Rightarrow (m_A^2 - m_1^2 - m_2^2 - 2p^2)^2 = 4(m_1^2 m_2^2 + (m_1^2 + m_2^2)p^2 + p^4)$$

$$\begin{aligned} \Rightarrow m_A^4 + m_1^4 + m_2^4 + 4p^4 - 2m_A^2(m_1^2 + m_2^2) + 2m_1^2 m_2^2 - 4p^2(m_A^2 - m_1^2 - m_2^2) \\ = 4m_1^2 m_2^2 + 4p^2(m_1^2 + m_2^2) + 4p^4 \end{aligned}$$

$$\Rightarrow p^2 = \frac{1}{4m_A^2} (m_A^4 + m_1^4 + m_2^4 - 2m_A^2(m_1^2 + m_2^2) - 2m_1^2 m_2^2)$$

Thus, in this case final result is

$$\Gamma_{A \rightarrow 1,2} = \frac{|M|^2}{16\pi m_A^3} (m_A^4 + m_1^4 + m_2^4 - 2m_A^2(m_1^2 + m_2^2) - 2m_1^2 m_2^2)^{1/2}$$

If $m_1 = m_2 = 0$, this simplifies to

$$\Gamma = \frac{|M|^2}{16\pi m_A}$$

6 Scattering

- Elastic scattering: $A+B \rightarrow A+B$

- Inelastic scattering: others,



.....

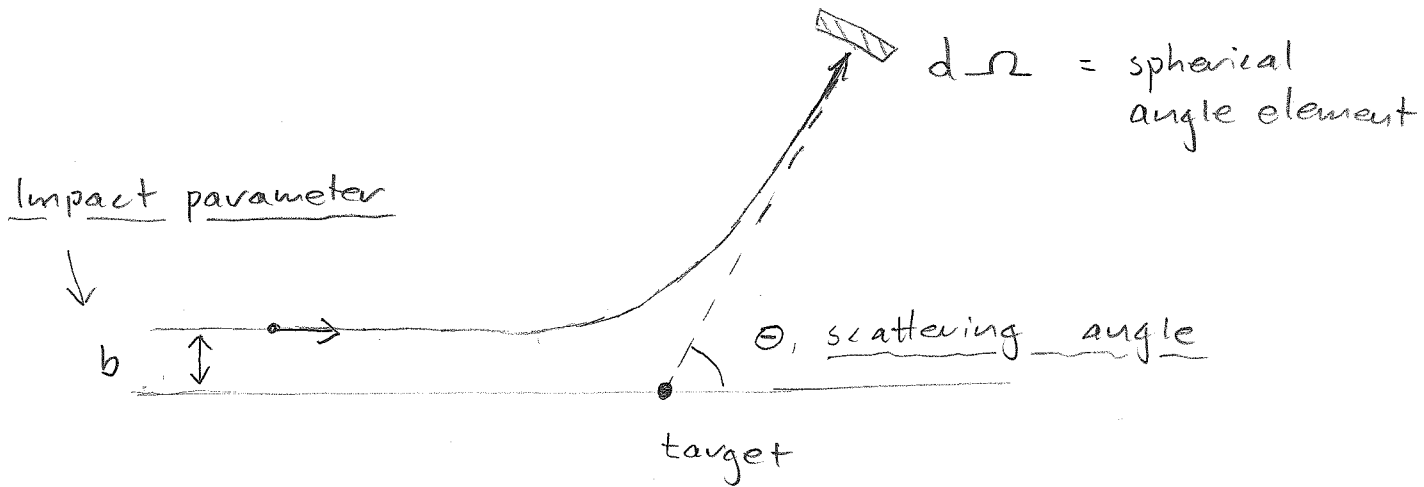
- Cross-section σ

- * Classically, cross-sectional area of the target:



- * Quantum mechanics: σ can depend dramatically on the energy of the collision. At resonance energy looks exceptionally large!

- Exclusive process: all products identified
- Inclusive process: only part identified; e.g.
 $A + B \rightarrow C + (\text{anything})$



$$d\Omega = \sin\theta d\theta d\phi \quad ; \quad \int d\Omega = 4\pi$$

- Differential cross-section $\frac{d\sigma}{d\Omega}$: scattering cross-section into $d\Omega$, to direction (θ, ϕ)

$$\sigma = \int d\Omega \frac{d\sigma}{d\Omega}$$

- Luminosity $L_{in} \equiv \frac{d^2 N_{in}}{dt dA}$: incoming particle density; particles/time/area unit

Now

$$\frac{dN_{out}}{dt d\Omega} = L_{in} \frac{d\sigma}{d\Omega}$$

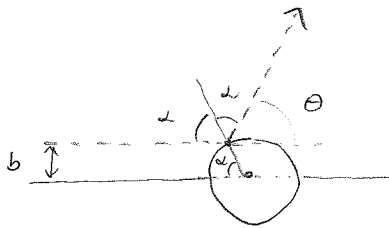
↑ particles/time/space angle to direction θ, ϕ

$$[L_{in}] = \frac{1}{m^2 s} \quad ; \quad [\sigma] = \left[\frac{d\sigma}{d\Omega} \right] = m^2 \sim \text{area}$$

- Common unit for σ : 1 barn = $(10 \text{ fm})^2$
= 10^{-24} cm^2

(Name comes from the fact that in particle physics this is very large, easy to hit "barn door".)

Example: classical hard sphere scattering
(point particle + hard sphere)



$$b = R \sin \alpha$$

$$2\alpha + \theta = \pi$$

$$\Rightarrow b = R \sin\left(\frac{\pi}{2} - \frac{\theta}{2}\right) = R \cos \frac{\theta}{2}$$

$$\text{Luminosity } L = \frac{d^2 N_{in}}{dt dA}$$

- Polar angle ϕ constant

- How many particles scatter to angle $(\theta, \theta + d\theta)$?

* All which have impact parameter between $b, b + db$, where

$$db = -R \sin \frac{\theta}{2} \frac{d\theta}{2}$$

$$d\Omega = d\phi \sin \theta d\theta$$

$$\text{Now } dA = |db| \cdot b d\phi$$

$$= \frac{R \sin \frac{\theta}{2}}{2} b d\theta d\phi$$

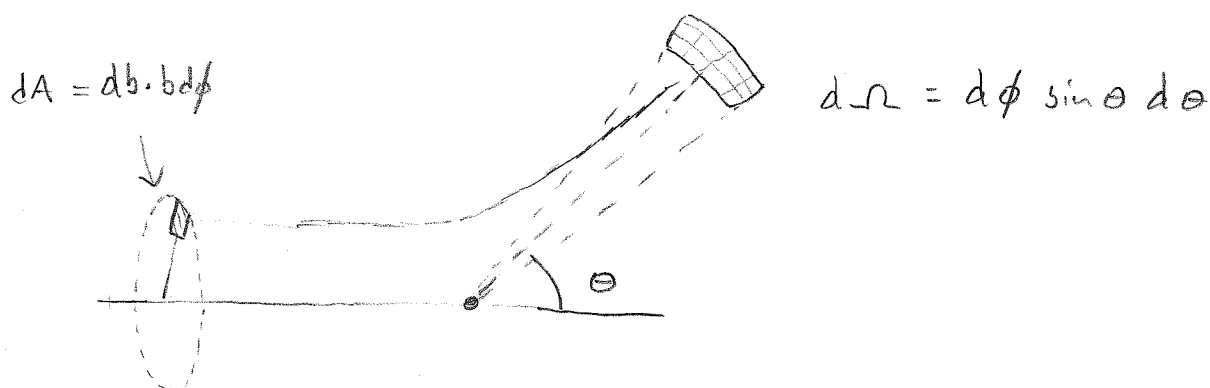
$$= \frac{R^2}{2} \sin \frac{\theta}{2} \cos \frac{\theta}{2} d\theta d\phi = \frac{R^2}{4} \sin \theta d\theta d\phi$$

Thus,

$$\frac{\frac{d^2 N}{dt d\Omega}}{\frac{d^2 N}{dt d\phi d\theta \sin\theta}} = \frac{\frac{d^2 N}{dt dA}}{\underbrace{\frac{R^2/4 \sin\theta}{\sin\theta}}_{L_{in}}} = L_{in} \cdot \frac{R^2}{4}$$

$$\Rightarrow \frac{d\sigma}{d\Omega} = \frac{R^2}{4} \Rightarrow \sigma = \int d\Omega \frac{R^2}{4} = \underline{\pi R^2}$$

The classical x-section! (of course...)



- All particles which go through area dA (in time dt) will exit through spherical angle $d\Omega$. To get $d\sigma/d\Omega$, we need to do change of variables \rightarrow |jacobian|.

• Note: Luminosity $L_{in} = \frac{d^2 N_{in}}{dt dA} =$ particles going through dA !

$$\Rightarrow L_{in} = \underbrace{\frac{N_{in}}{V}}_{\text{density of incoming particles}} v_{in} = \frac{1}{V} v_{in}$$

density of incoming particles

↑ if we normalise

$\langle A|A \rangle = 1$, 1 particle/ V !